

Study of Thermal Stability in A Newtonian Fluids

Satya Ranjan Pradhani¹, Dr. K. B. Singh², Dr. L. K. Roy³

¹Research Scholar, University Department of Mathematics, B. R. A. Bihar University, Muzaffarpur, Bihar, India

²P. G. Department of Physics, L. S. College, B. R. A. Bihar University, Muzaffarpur, Bihar, India

³Department of Mathematics, T. P. Verma College, Narkatiaganj, West Champaran, B. R. A. Bihar University, Muzaffarpur, Bihar, India

ABSTRACT

Article Info

Volume 9, Issue 5

Page Number : 111-118

Publication Issue

September-October-2022

Article History

Accepted : 05 Sep 2022

Published : 18 Sep 2022

In this present paper, we studied about thermal stability in a Newtonian fluid. A detailed account of thermal convection in a Newtonian fluid layer in the presence of magnetic field has been given by Chandrasekhar [11]. Stokes [79] has formulated the theory of couple-stresses in fluids. The theory due to Stokes allows for polar effects such as the presence of couple stresses and body couples. The theory has been applied to the study of some simple lubrication problems. According to the theory, couple-stresses are found to appear in noticeable magnitudes in fluids with very large molecules. Since the long chain hyaluronic acid molecules are found as additives in synovial fluid, Walicki and Walicka [80] modeled synovial fluid as a couple-stress fluid in human joints. A human joint is a dynamically loaded bearing which has particular cartilage as the bearing and synovial fluid as the lubricant. A synovial fluid is the natural lubricant of joints of the vertebrates. The synovial is obtained from the hyaluronic acid, a fluid of high viscosity, near a gel. Normal synovial fluid is clear or yellowish and is a non-Newtonian, viscous fluid. The shoulder, hip, knee and ankle joints are loaded with the synovial fluids and joints have a low friction coefficient and negligible wear.

Keywords: Thermal Stability, Newtonian Fluid, Instability, Rotation, Magnetic Field.

I. INTRODUCTION

The presence of small amounts of additives in a lubricant can improve bearing performance by increasing the lubricant viscosity and thus, producing the increase in the load capacity. These additives in a lubricant also reduce the coefficient of friction and increase the temperature range in which the bearing

can operate. Lin [51] showed that the bearing with couple stresses in fluid as the lubricant improves the squeeze film characteristic and results in a longer bearing life. All diseases of joints are caused by or connected with a malfunction of the lubrication. One of the applications of couple-stresses in fluid is its use in the study of the mechanism of lubrication of synovial joints, which has become the objective of

scientific research. The problem of couple-stress fluid heated from below in porous medium is considered by Sharma and Sharma [225] and Kumar [226]. Sharma and Aggarwal [203] and Singh et al [227] considered the effect of compressibility, suspended particles and rotation on thermal convection in an elastic-viscous fluid in hydromagnetic.

and the layer is acted upon by the gravity field $g(0,0,-g)$, a uniform vertical magnetic field $H(0,0,H)$ and rotation $\Omega(0,0,\Omega)$.

II. MATHEMATICAL FORMULATION

Consider an infinite, horizontal, electrically conducting, incompressible, couple-stress fluid layer to thickness d , bounded by the planes $z=0$ and $z=d$. This fluid layer is heated from below so that a uniform temperature gradient $\beta \left(\left| \frac{dT}{dz} \right| \right)$ is maintained

Let $p, \rho, T, a, v, \mu', k, \text{ and } q(u, v, w)$ denote respectively pressure, density, temperature, thermal coefficient of expansion, kinematic viscosity, couple-stress viscosity, thermal diffusivity and velocity of the fluid. The equation of motion, continuity and heat conduction of couple-stress fluid are

$$\frac{\partial q}{\partial t} + (q \cdot \nabla)q = -\frac{1}{\rho_0} \nabla p + g \left(1 + \frac{\delta \rho}{\rho_0} \right) + \left(v - \frac{\mu'}{\rho_0} \nabla^2 \right) \nabla^2 q + 2(q \times \Omega) + \frac{\mu_e}{4\pi\rho_0} [\nabla \times H \times H] \quad (1)$$

$$\nabla \cdot q = 0 \quad (2)$$

$$\frac{\partial T}{\partial t} + (q \cdot \nabla)T = k \nabla^2 T \quad (3)$$

$$\frac{\partial H}{\partial t} = (H \cdot \nabla)q + \eta \nabla^2 H \quad (4)$$

$$\nabla \cdot H = 0 \quad (5)$$

The equation of state is

$$\rho = \rho_0 [1 - \alpha(T - T_0)] \quad (6)$$

where the suffix zero refers to the values at the reference level $z=0$.

The basic motionless solution is

$$q = (0,0,0), p = p(z), T = T_0 - \beta z, \rho = \rho_0(1 + \alpha\beta z), N = N_0, \text{ a constant.} \quad (7)$$

Assume small perturbations around the basic solution and let $\delta p, \delta \rho, \theta, N, H (h_x, h_y, H + h_z)$ and $q(u,v,w)$ denote respectively the perturbations in pressure, density, temperature, number density, magnetic field and couple-stress fluid velocity $(0,0,0)$. The change in density $\delta \rho$ caused mainly by the perturbation θ in temperature is given by

$$\delta \rho = -\alpha \rho_0 \theta \quad (8)$$

Then the linearized perturbation equations of the couple-stress fluid become

$$\frac{\partial q}{\partial t} = -\frac{1}{\rho_0} \nabla \delta p - g \alpha \theta + \left(v - \frac{\mu'}{\rho_0} \nabla^2 \right) \nabla^2 q + 2(q \times \Omega) + \frac{\mu_e}{4\pi\rho_0} [\nabla \times H \times H] \quad (9)$$

$$\nabla \cdot q = 0 \quad (10)$$

$$\frac{\partial \theta}{\partial t} = \beta w + k \nabla^2 \theta \quad (11)$$

$$\frac{\partial q}{\partial t} = (H \cdot \nabla)q + \eta \nabla^2 h \quad (12)$$

$$\nabla \cdot h = 0 \quad (13)$$

where $k = \frac{q}{\rho_0 c_v}$

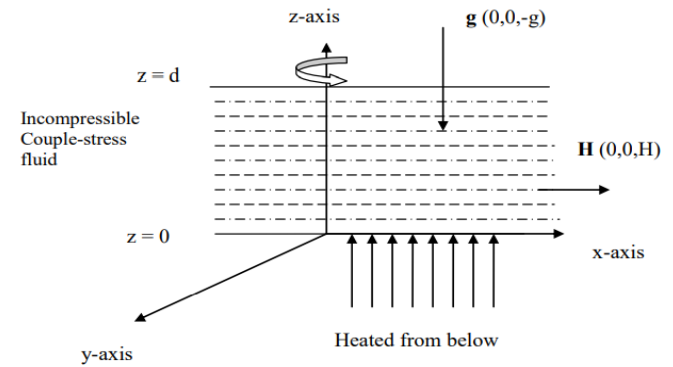


Figure 1: Geometrical configuration

2.1 Dispersion Relation

Analyze the perturbations in to normal modes by seeking solutions in the form $[w, \theta, h, \zeta, \xi] = [W(z), \Theta(z), K(z), Z(z), X(z)] \exp(ik_x x + ik_y y + nt)$. where k_x, k_y are the wave numbers along x and y directions respectively, and $k = (k_x^2 + k_y^2)^{1/2}$ is the resultant wave number of the disturbance and n is the growth rate which is, in general, a complex constant and $\zeta = \frac{\partial v}{\partial x} - \frac{\partial u}{\partial y}$ and $\xi = \frac{\partial h_y}{\partial x} - \frac{\partial h_x}{\partial y}$ stand for the z-components of vorticity and current density, respectively.

The non-dimensional system of equations eliminating the physical quantities is

$$[\sigma - (1 - F(D^2 - a^2))(D^2 - a^2)](D^2 - a^2)W = -\frac{g\alpha a^2 d^2}{v}\Theta - \frac{2\Omega d^3}{v}DZ + \frac{\mu_e H d}{4\pi\rho_0 v}(D^2 - a^2)DK \tag{14}$$

$$[\sigma - (1 - F(D^2 - a^2))(D^2 - a^2)]Z = \left(\frac{2\Omega d}{v}\right)DW + \frac{\mu_e H d}{4\pi\rho_0 v}DK \tag{15}$$

$$(D^2 - a^2 - \sigma p_1)\Theta = -\left(\frac{\beta d^2}{k}\right)W \tag{16}$$

$$(D^2 - a^2 - p_2\sigma)K = -\frac{Hd}{\eta}DW \tag{17}$$

$$(D^2 - a^2 - p_2\sigma)X = -\frac{Hd}{\eta}DZ \tag{18}$$

Eliminating Z, X, Θ and K between equations (14) – (18), we obtain

$$[\sigma + F(D^2 - a^2)^2 - (D^2 - a^2)][D^2 - a^2 - \sigma p_1][D^2 - a^2 - \sigma p_2][D^2 - a^2]W + R\lambda a^2[D^2 - a^2 - \sigma p_2]W + Q[D^2 - a^2 - \sigma p_1][D^2 - a^2]D^2W \tag{19} + T_A \frac{[D^2 - a^2 - \sigma p_2]^2 [D^2 - a^2 - \sigma p_1] D^2 W}{\{[\sigma + F(D^2 - a^2) - (D^2 - a^2)](D^2 - a^2 - \sigma p_2) + QD\}} = 0$$

where $R = \frac{g\alpha\beta d^4}{vk}$ is the thermal Rayleigh number, $T_A = \left(\frac{2\Omega d^2}{y}\right)^2$ is the Taylor's number and $Q = \frac{\mu_e H^2 d^2}{4\pi\rho_0 v\eta}$ is the Chandrasekhar number.

Consider the case in which both the boundaries are free, the medium adjoining the fluid is perfectly conducting and temperatures at the boundaries are kept fixed. The boundary conditions, appropriate for the problem, are $W = 0 = Z = \Theta$ and $D^2W = 0$ at $z = 0$ and $z = 1$ (20)

$$W = W_0 \sin \pi z \tag{21}$$

where W_0 is constant. Substituting the proper solution (21) to equation (19), we obtain the dispersion relation

$$R_1 = \frac{(1+x)}{\lambda x} [i\sigma + F_1 + (1+x)][1+x+i\sigma p_1] + \frac{Q_1(1+x)[1+x+i\sigma_1 p_1]}{\lambda x[1+x+i\sigma_1 p_1]} + \frac{T_{A_1}[1+x+i\sigma_1 p_2][1+x+i\sigma_1 p_1]}{\lambda x\{[i\sigma + F_1(1+x)^2 + (1+x)](1+x+i\sigma_1 p_2) + Q_1\}} \tag{22}$$

where, $R_1 = \frac{R}{\pi^4}$, $T_{A_1} = \frac{T_A}{\pi^4}$, $i\sigma_1 = \frac{\sigma}{\pi^2}$, $Q_1 = \frac{Q}{\pi^2}$ and $F_1 = \pi^2 F$

The above relation expresses the modified Rayleigh number R_1 as a function of couple stress parameter F_1 , rotation parameter T_{A_1} , magnetic field parameter Q_1 and dimension less wave number x.

2.2 The Stationary Convection

For stationary convection, the marginal state will be characterized by $\sigma = 0$. Thus equation (22) reduces to

$$R_1 = \frac{(1+x)}{\lambda x} \left[\{F_1(1+x) + 1\}(1+x)^2 + Q_1 + \frac{T_{A_1}(1+x)}{\{[F_1(1+x)+1\](1+x)^2 + Q_1\}} \right] \tag{23}$$

To study the effect of suspended particles, rotation, couple-stress and magnetic field, we examine the nature of $\frac{dR_1}{dT_{A_1}}$, $\frac{dR_1}{dF_1}$ and $\frac{dR_1}{dQ_1}$ analytically. Equation (23) gives

$$\frac{dR_1}{dT_{A_1}} = \left(\frac{1+x}{x}\right) \left\{ \frac{1+x}{[1+F_1(1+x)](1+x)^2 + Q_1} \right\} \tag{24}$$

which shows that rotation has a stabilizing effect on the system.

Also, from equation (3.1.23), we have

$$\frac{dR_1}{dF_1} = \frac{(1+x)^4}{x} \left\{ 1 - T_{A_1} \frac{1+x}{\{[1+F_1(1+x)](1+x)^2 + Q_1\}^2} \right\} \tag{25}$$

$$\frac{dR_1}{dQ_1} = \left(\frac{1+x}{x}\right) \cdot \left\{ T_{A_1} \frac{1+x}{\{[1+F_1(1+x)](1+x)^2+Q_1\}^2} + 1 \right\} \tag{26}$$

which shows that couple-stresses and magnetic field have a stabilizing or destabilizing effect on the system according to $T_{A_1}(1+x) < or > \{[1+F_1(1+x)](1+x)^2+Q_1\}^2$

In the absence of rotation ($T_{A_1} = 0$), we have

$$\frac{dR_1}{dF_1} = \frac{(1+x)^4}{x} \tag{27}$$

$$\frac{dR_1}{dQ_1} = \left(\frac{1+x}{x}\right) \tag{28}$$

which shows that couple-stresses and magnetic field clearly have a stabilizing effect on the system.

In the absence of magnetic field ($Q_1 = 0$), we have

$$\frac{dR_1}{dF_1} = \frac{(1+x)^4}{x} \cdot \frac{1}{H} \left\{ 1 - T_{A_1} \frac{1}{[1+F_1(1+x)]^2(1+x)^3} \right\} \tag{29}$$

which shows that couple-stress has a stabilizing (destabilizing) effect on the system according as $T_{A_1} < or > (1+x)^3[1+F_1(1+x)]^2$

The dispersion relation (22) is also analyzed numerically. In figure 2, R_1 is plotted against x for $T_{(A_1)}=100,150,200, F_1=0.5$ and $Q_1=10$. In figure 3, R_1 is plotted against notation parameter $T_{(A_1)}$ for various values of wave number x . In both the figures, it is found that rotation postpones the onset of convection as the Rayleigh number increases with the increase in rotation parameter. In figure 4, R_1 is plotted against x for $Q_1=20,40,60; F_1=0.2, T_{(A_1)}=70$ and in figure 5 R_1 is plotted against Q_1 for $x =12,4,8,10$. Here it is observed that the magnetic field hastens the onset of convection for small wave numbers as the Rayleigh number decreases with an increase in the magnetic field parameter and postpones the onset of convection for higher wave numbers as the Rayleigh number decreases with the increase in couple-stress parameter and postpones the onset of convection for higher wave numbers as the Rayleigh number increases with the increase in couple-stress parameter.

The critical Rayleigh numbers listed in tables 1 to 3 and illustrated in figures 8-10 are obtained from figures 2 to 7 by locating the minimum numerically. From table 1 and figure 8, it is clear that rotation has stabilizing effect on the system. From the table 2 and figure 9, it is observed that the magnetic field has stabilizing effect in the absence of rotation, destabilizing effect for $T_{(A_1)}=6000$ and for $T_{(A_1)}=2000$, the value of critical Rayleigh number

first decreases and then increases for the increase in the value of magnetic field parameter. In figure 10, critical Rayleigh number R_{cis} plotted against couple-stress parameter F_1 . In the absence of rotation, couple-stresses have stabilizing effect whereas in the presence of rotation, the value of critical Rayleigh number R_c decreases and then increases for the increase in value of couple-stress parameter F_1 . Table 3 confirms these results numerically.

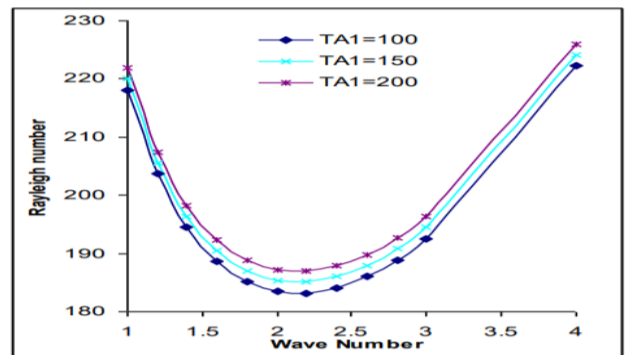


Figure 2: Variation of R_1 with x

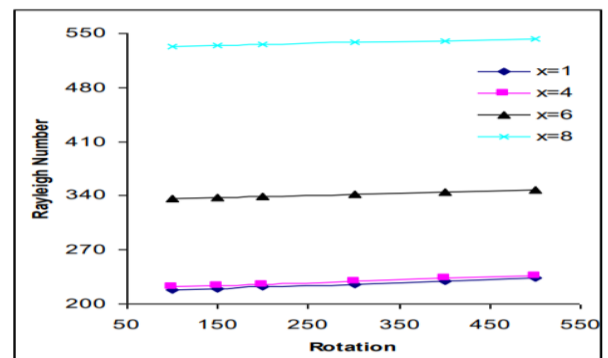


Figure 3: Variation of R_1 with T_{A1}

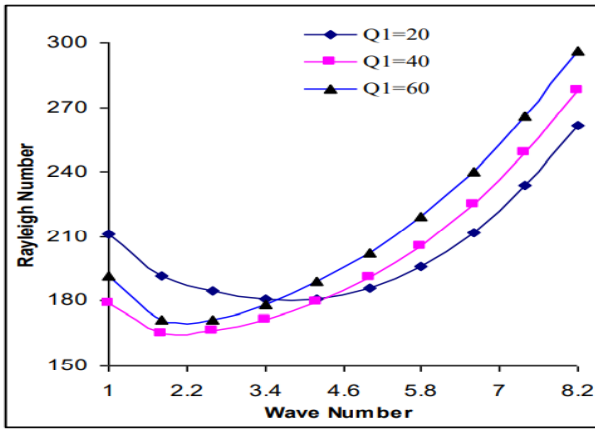


Figure 4: Variation of R1 with x

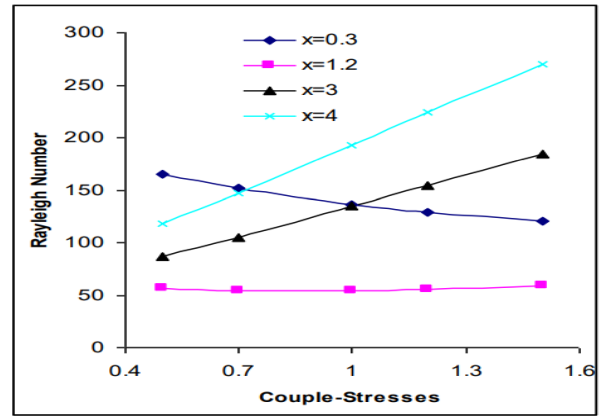


Figure 7: Variation of R1 with F1

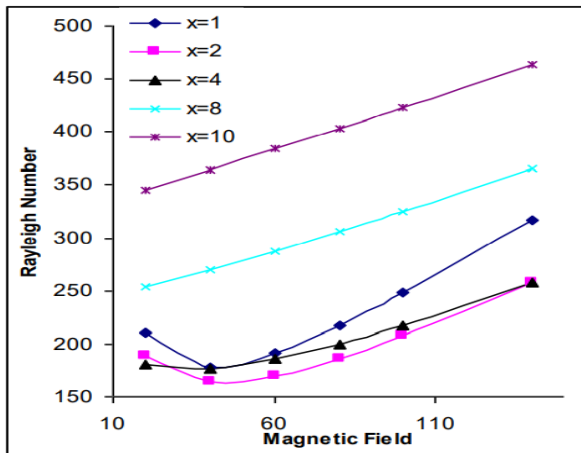


Figure 5: Variation of R1 with Q1

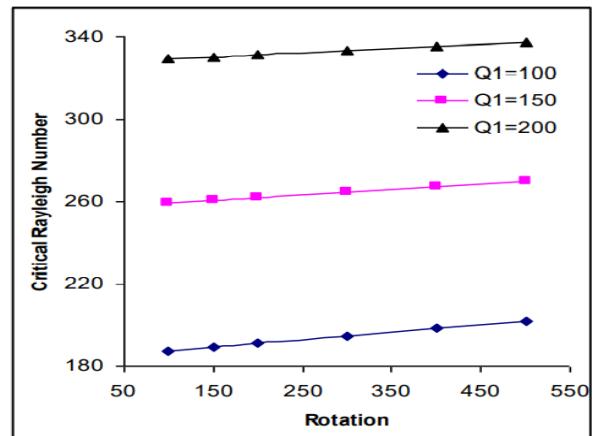


Figure 8: Variation of Rc with TA1

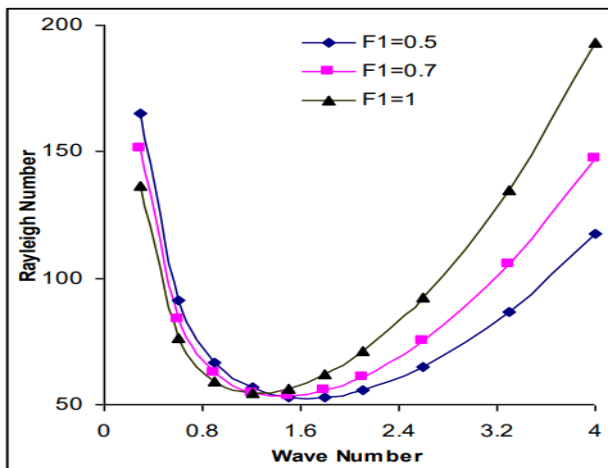


Figure 6: Variation of R1 with x

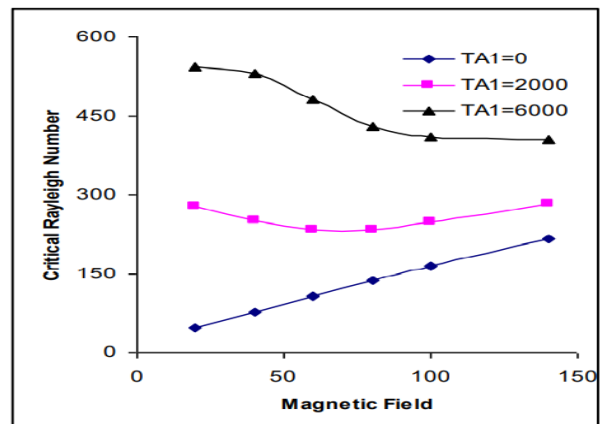


Figure 9: Variation of Rc with Q1

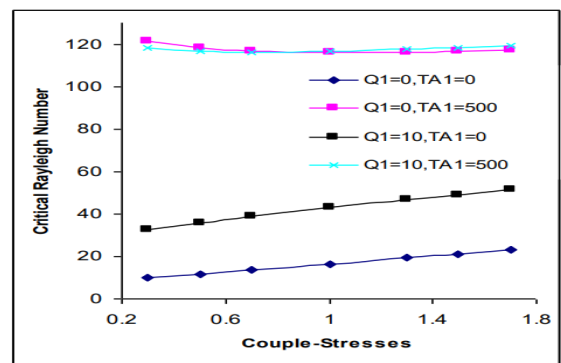


Figure 10: Variation of Rc with F1

Table 1: The critical Rayleigh numbers and the wave numbers of the associated disturbances for the onset of instability as stationary convection for various values of T_{AI} .

T_{AI}	$Q_1 = 100$		$Q_1 = 150$		$Q_1 = 200$	
	x_c	R_c	x_c	R_c	x_c	R_c
100	2.0	187.4235	2.3	259.2703	2.5	329.2593
150	2.0	189.2602	2.3	260.5939	2.5	330.3077
200	2.0	191.0969	2.3	261.9175	2.5	331.3561
250	2.0	194.7704	2.3	264.5647	2.5	333.4530
300	2.0	198.4439	2.3	267.2120	2.5	335.5498
400	2.0	202.1173	2.3	269.8592	2.5	337.6466

Table 2: The critical Rayleigh numbers and the wave numbers of the associated disturbances for the onset of instability as stationary convection for various values of Q_1 .

Q_1	$T_{AI} = 0$		$T_{AI} = 2000$		$T_{AI} = 6000$	
	x_c	R_c	x_c	R_c	x_c	R_c
20	1.6	46.3411	5.5	277.2780	8.4	543.9253
40	2.0	77.5500	2.0	251.6312	7.5	529.6470
60	2.3	106.8679	2.0	233.0730	1.6	481.3190
80	2.5	135.1525	2.1	233.5286	1.7	430.0211
100	2.7	162.7387	2.3	246.9746	1.8	408.6656
140	3.1	216.5092	2.7	281.4301	2.1	405.1923

Table 3: Critical Rayleigh numbers and the wave numbers of the associated disturbances for the onset of instability as stationary convection for various values of F_1 .

F_1	$Q_1 = 0$				$Q_1 = 10$			
	$T_{AI} = 0$		$T_{AI} = 500$		$T_{AI} = 0$		$T_{AI} = 500$	
	x_c	R_c	x_c	R_c	x_c	R_c	x_c	R_c
0.3	0.4	9.7412	3.3	121.3369	1.0	32.8000	2.7	118.4836
0.5	0.4	11.6620	2.8	118.4081	0.9	35.9723	2.3	116.6769
0.7	0.4	13.5828	2.4	116.9902	0.9	38.8683	2.0	116.2168
1.0	0.4	16.4640	2.1	116.2354	0.8	42.9120	1.9	116.6277
1.3	0.4	19.3452	1.9	116.3986	0.7	46.8153	1.6	117.6984
1.7	0.4	23.1868	1.6	117.1955	0.7	51.5879	1.4	119.7011

III. STABILITY OF THE SYSTEM AND OSCILLATORY MODES

Now to determine under what conditions the principle of exchange of stabilities (PES) is satisfied (i.e. is real σ is real and the marginal states are characterized by $\sigma=0$) and the oscillations come into play, we multiply equation (14) with W^* and integrate over the range of z and making use of equations (15)-(18) together with the boundary conditions (20) and get

$$\sigma I_1 + I_2 + F I_3 - \frac{g a k a^2}{\nu \beta} (I_4 + H \sigma^* p_1 I_5) + d^2 [\sigma I_6 + I_7 + F I_8] + \frac{\mu_e \eta d^2}{4 \pi \rho_o \nu} (I_9 + \sigma^* p_2 I_{10}) + \frac{\mu_e \eta}{4 \pi \rho_o \nu} (I_{12} + \sigma^* p_2 I_{11}) = 0$$

where,

$$\begin{aligned}
 I_1 &= \int (|DW|^2 + a^2|W|^2)dz, & I_2 &= \int (|D^2W|^2 + 2a^2|DW|^2 + a^4|W|^2)dz, \\
 I_3 &= \left(\int |D^2W|^2 + 3a^2|D^2W|^2 + 3a^4|DW|^2 + a^6|W|^2 \right) dz, \\
 I_4 &= \int (|D\Theta|^2 + a^2|\Theta|^2) dz, & I_5 &= \int |\Theta|^2 dz, \\
 I_6 &= \int |Z|^2 dz & I_7 &= \int (|DZ|^2 + a^2|Z|^2) dz, \\
 I_8 &= \int (|D^2Z|^2 + 2a^2|DZ|^2 + a^4|Z|^2) dz, \\
 I_9 &= \int (|DX|^2 + a^2|X|^2)dz, & I_{10} &= \int |X|^2 dz, \\
 I_{11} &= \int (|DK|^2 + a^2|K|^2)dz, & I_{12} &= \int (|D^2K|^2 + 2a^2|DK|^2 + a^4|X|^2)dz.
 \end{aligned}$$

and σ^* is complex conjugate of σ . The integrals $I_1 - I_{12}$ are all positive definite. Putting $\sigma = i\sigma_i (\sigma^* = -i\sigma_i)$ in equation (30) and equating imaginary parts, we obtain

$$\sigma_i \left[I_1 + \frac{gaka^2}{v\beta} (I_4 + Hp_1I_5) + d^2I_6 - \frac{\mu_e\eta d^2}{4\pi\rho_0v} p_2I_{10} - \frac{\mu_e\eta}{4\pi\rho_0v} p_2I_{11} \right] = 0 \tag{31}$$

It is clear from equation (31) that σ_i may be zero or non-zero, which implies that modes may be non-oscillatory or oscillatory. In the absence of magnetic field and rotation, equation (31) reduces to

$$\sigma_i \left[I_1 + \frac{gaka^2}{v\beta} (I_4 + Hp_1I_5) + d^2I_6 \right] = 0 \tag{32}$$

The terms in the bracket are positive definite. Thus $\sigma_i = 0$ which means that the oscillatory modes are not allowed and the principle of exchange of stabilities is satisfied in the absence of rotation and magnetic field.

IV. CONCLUSION

In this section, the effect of magnetic field and rotation has been considered on the thermal stability of a couple-stress fluid. The effect of various parameters such as magnetic field, rotation and couple-stresses has been investigated analytically as well as numerically. The main results from the analysis are as follows: In order to investigate the effects of magnetic field, rotation and couple-stresses, we examine the behavior of $\frac{dR_1}{dQ_1}, \frac{dR_1}{dT_{A_1}}$ and $\frac{dR_1}{dF_1}$ analytically. It is found that rotation has stabilizing effect on the system. The magnetic field couple-stresses has a stabilizing effect in the absence of rotation whereas in the presence of rotation it has a stabilizing effect if $T_{A_1}(I+x) < \{[I + F_1(I+x)](I+x)^2 + Q_1\}^2$ and destabilizing effect if $T_{A_1}(I+x) < \{[I + F_1(I+x)](I+x)^2 + Q_1\}^2$. The principle of exchange of stabilities is satisfied in the absence of rotation and magnetic field.

V. REFERENCES

- [1]. Yuan S. W., “Foundations of Fluid Mechanics”, Prentice Hall of India, New Delhi, 1988.
- [2]. Alfvén H., “On the Existence of Electromagnetic Hydromagnetic Waves”, Ark. Mat. Astr. Fys., vol. 29, no.1, 1942.
- [3]. Batchelor G. K., “On the Spontaneous Magnetic Field in a Conducting Liquid in Turbulent Motion”, Proc. Roy. Soc. London, vol. 210(A), pp. 405-416, 1950.
- [4]. Lundquist S., “On Stability of Magnetohydrostatic Fluids”, Phys. Rev., vol. 83(2), pp. 307- 311, 1951.
- [5]. Lehnert B., “Magneto Hydrodynamic Waves in Liquids Sodium”, Phys. Rev., vol.94, pp. 815-824, 1954.

- [6]. Sutton G. W., Sherman A., "Engineering Magnetohydrodynamics", McGraw Hill, New York, 1965.
- [7]. Roberts P. H., "Introduction to Magnetohydrodynamics", Longmans Green, London, 1967.
- [8]. Cowling T. G., "Magnetohydrodynamics", 2nd ed., Adam Hilger, England, 1976.
- [9]. Bateman G., "MHD Instabilities", MIT press, Cambridge, 1978.
- [10]. Moffatt H. K., "Magnetic Field Generation in Electrically Conducting Fluids", Cambridge Univ. Press, England, 1978.
- [11]. Chandrasekhar S., "Hydrodynamic and Hydromagnetic Stability", Dover publications, New York, 1981.

Cite this article as :

Satya Ranjan Pradhani, Dr. K. B. Singh, Dr. L. K. Roy, "Study of Thermal Stability in A Newtonian Fluids ", International Journal of Scientific Research in Science and Technology (IJSRST), Online ISSN : 2395-602X, Print ISSN : 2395-6011, Volume 9 Issue 5, pp. 111-118, September-October 2022. Available at doi : <https://doi.org/10.32628/IJSRST229523>
Journal URL : <https://ijsrst.com/IJSRST229523>